

CP violation in leptogenesis and at low energy

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The problem :

Relation between low energy CP violation and CP violation in flavoured leptogenesis

Summary :

- Baryon asymmetry
- Thermal leptogenesis
- Effect of flavours in leptogenesis
- Sensitivity of flavoured leptogenesis from low energy CP-odd observables
- Conclusions

Cosmological matter excess

Observations :

- The known universe is made of matter (no γ rays from annihilations)
- ⇒ there is an excess of matter over anti-matter
- Matter $\simeq H = p + e^-$, implying a baryon asymmetry :

$$Y_B = \frac{n_B - n_{\bar{B}}}{n_\gamma} \neq 0$$

Measurements of baryon number density :

- Luminous matter : $Y_B \sim \text{few} \times 10^{-11}$
- BBN : $Y_B \sim 2 - 6 \times 10^{-10}$
- WMAP : $Y \sim 6.15 \pm .025 \times 10^{-10}$

No definite information on **lepton asymmetry** (there is an undetectable CMB of ν s which could contain a large lepton asymmetry)

If Universe was born with a baryon asymmetry
⇒ BA would have been diluted during *inflation*

Dynamical baryon asymmetry production after inflation :

- 1 **B violation** : to evolve from a state with $B = 0$ to $B \neq 0$
- 2 **C and CP violation** : different behaviour of particles and anti-particles
- 3 **out-of-equilibrium dynamics** : CPT conservation ⇒ particles and antiparticles with same mass ⇔ same abundance

In the Standard Model :

- 1 **B violation** from quantum anomalies and non perturbative processes (sphalerons)
- 2 **C and CP violation** (in the CKM matrix), but CP violation is too small
- 3 **out-of-equilibrium** in electroweak phase transition, but too smooth the shift of the Higgs vev for $m_h > 70$ GeV

⇒ **Extension of the Standard Model scenario**

$$\mathcal{L}_{\text{seesaw}} = (\bar{\ell}_L^i H_d^*) \mathbf{Y}_{\mathbf{e}}^*{}_{ij} e_R^j + (\bar{\ell}_L^i H_u^*) \lambda^*{}_{ij} N^j + \bar{N}^c{}^J \frac{\mathbf{M}^*{}_{JK}}{2} N^K + h.c.$$

At low energy ($\ll M_1$):

- Heavy degrees of freedom are integrated out
 - $[m_\nu] = \lambda M^{-1} \lambda^T v_u^2 = U D_\nu U^T$: effective light neutrino mass matrix
 - D_e : charged Yukawa couplings
- ⇒ 12 parameters in the lepton sector

Type 1 seesaw has 21 parameters (bottom-up parametrisation):

- D_ν : 3 light neutrino masses
- U_{MNS} : 3 mixing angles, Dirac phase δ , Majorana phases α and β
- $\lambda \lambda^\dagger = V_L^\dagger D_\lambda^2 V_L$: 3 neutrino Yukawa couplings, 3 mixing angles, 3 phases
- D_{Y_e} : 3 charged Yukawa couplings

Thermal leptogenesis

$$\mathcal{L}_{\text{seesaw}} = (\bar{\ell}_L^i H_d^*) \mathbf{Y}_{eij}^* e_R^j + (\bar{\ell}_L^i H_u^*) \lambda^*_{ij} N^j + \overline{N^c}^J \frac{M^*_{JK}}{2} N^K + h.c.$$

- Hierarchical N masses : $M_1 \sim 10^9 \text{ GeV} \gg M_2, M_3$
- Thermal production of the N_1 (and negligible production of N_2)

The process :

- N_1 produced by scattering processes at $T \sim M_1$
- N_1 decay violating CP \Rightarrow production of lepton asymmetry
- If inverse decays are out of equilibrium the asymmetries may survive
- The lepton asymmetry is converted into baryon asymmetry by sphalerons

Thermal leptogenesis

- Produce the maximal thermal N density if production rate is fast enough (e.g. $\Gamma(q_L t_R \rightarrow \phi \rightarrow \ell N)$) :

$$\Gamma_{prod} \sim \frac{h_t^2 \lambda^2}{4\pi} T > H$$

- Lepton asymmetry produced if N decay are CP violating and out of equilibrium and inverse decays are turn off :

$$\epsilon = \frac{\Gamma(N_1 \rightarrow \phi \ell) - \Gamma(N_1 \rightarrow \bar{\phi} \bar{\ell})}{\Gamma(N_1 \rightarrow \phi \ell) + \Gamma(N_1 \rightarrow \bar{\phi} \bar{\ell})}$$

$$\Gamma_{ID} \equiv \Gamma(I\phi \rightarrow N_1) = \Gamma_{decay} e^{-M_1/T} = \frac{[\lambda\lambda^\dagger]_{11} M_1}{8\pi} e^{-M_1/T} < H = \frac{10T^2}{m_{pl}}$$

- When $\Gamma_{ID} \sim H \Rightarrow n_N/n_\gamma \simeq e^{-M_1/T} \simeq H/\Gamma_{decay}$
- Lepton asymmetry \rightarrow baryon asymmetry :

$$Y_B \sim \frac{1}{3} Y_L = \frac{n_l - n_{\bar{l}}}{n_N} \frac{n_N}{n_\gamma} \sim \frac{H}{3g_*} \frac{\epsilon}{\Gamma_{decay}}$$

Flavour in leptogenesis

- Rates for interactions involving charged lepton yukawas :

$$\Gamma_\alpha \sim 5 \times 10^{-3} h_\alpha^2 T$$

If this rates are in equilibrium \Rightarrow flavours become distinguishable¹²³ :

- $\Gamma_\tau > H$ for $T < 10^{12}$ GeV
 - $\Gamma_\mu > H$ for $T < 10^9$ GeV
- \Rightarrow Between 10^9 GeV and 10^{12} GeV we can distinguish the τ flavour (which is in equilibrium) from the others :

$$\hat{\ell}_o = \frac{\lambda_{\mu 1} \hat{\mu} + \lambda_{e 1} \hat{e}}{\sqrt{|\lambda_{\mu 1}|^2 + |\lambda_{e 1}|^2}}$$

- The lepton asymmetries ϵ_τ and ϵ_0 evolve **separately**

¹R. Barbieri, P. Creminelli, A. Strumia and N. Tetradis, hep-ph/9911315

²A. Abada, S. Davidson, F. X. Josse-Michaux, M. Losada and A. Riotto, hep-ph/0601083

³E. Nardi, Y. Nir, E. Roulet and J. Racker, hep-ph/0601084

Flavour in leptogenesis

- Produce the maximal thermal N density if production rate is fast enough (e.g. $\Gamma(q_L t_R \rightarrow \phi \rightarrow \ell_\alpha N)$) :

$$\Gamma_{prod} = \sum_\alpha \Gamma_{prod}(\rightarrow \phi \ell_\alpha) \sim \sum_\alpha \frac{h_t^2 |\lambda_{\alpha 1}^2|}{4\pi} T > H$$

- Lepton asymmetry in flavour α (produced in N decay) survives when inverse decays from flavour α are turn off :

$$\epsilon_{\alpha\alpha} = \frac{\Gamma(N_1 \rightarrow \phi \ell_\alpha) - \Gamma(N_1 \rightarrow \bar{\phi} \bar{\ell}_\alpha)}{\Gamma(N_1 \rightarrow \phi \ell_\alpha) + \Gamma(N_1 \rightarrow \bar{\phi} \bar{\ell}_\alpha)}$$

$$\Gamma_{\alpha\alpha} \equiv \Gamma(l_\alpha \phi \rightarrow N_1) \simeq \frac{|\lambda_{\alpha 1}^2| M_1}{8\pi} e^{-M_1/T} < H = \frac{10T^2}{m_{pl}}$$

- At T_α , when inverse decays from flavour α turn off :

$$\frac{n_N}{n_\gamma}(T_\alpha) \simeq e^{-M_1/T_\alpha} \simeq \frac{H}{\Gamma_{\alpha\alpha}}$$

$$Y_B \sim \frac{1}{3} \sum_\alpha \frac{n_l - n_{\bar{l}}}{n_N} \frac{n_N}{n_\gamma} \sim \frac{H}{3g_*} \sum_\alpha \frac{\epsilon_{\alpha\alpha}}{\Gamma_{\alpha\alpha}} \neq \frac{H}{3g_*} \frac{\sum_\alpha \epsilon_{\alpha\alpha}}{\sum_\beta \Gamma_{\beta\beta}}$$

The question

- The relation between CP violation at low energy (measurable in neutrino oscillations) and CP violation at high energy ?

Given the measured value of the baryon asymmetry, can an allowed range for the U_{MNS} phases be predicted ?

- First answer in Branco et al.⁴ in leptogenesis without flavour
- Easily seen with Casas-Ibarra parametrisation and hierarchical RH neutrinos :

$$\lambda = UD_k^{1/2} RD_M^{1/2} \quad \Rightarrow \quad \epsilon = -\frac{3M_1}{16\pi v^2} \frac{\Im(\sum_\rho m_\rho^2 R_{\rho 1}^2)}{\sum_\beta m_\beta |R_{1\beta}|^2}$$

$\Rightarrow \epsilon$ is independent of U_{MNS} phases

- We want to address the same problem in **flavoured leptogenesis**

⁴Branco, Morozumi, Nobre and Rebelo, hep-ph/0107164

Analytic result

In flavoured leptogenesis :

$$\epsilon_{\alpha\alpha} = -\frac{3M_1}{16\pi v^2} \frac{\Im(\sum_{\beta\rho} m_\beta^{1/2} m_\rho^{3/2} U_{\alpha\beta}^* U_{\alpha\rho} R_{\beta 1} R_{\rho 1})}{\sum_{\beta} m_\beta |R_{1\beta}|^2}$$

We look for an area of the parameter space where :

- We have enough baryon asymmetry
- Y_B is independent from low energy phases

It is found :

- In strong wash-out regime
- Using a simple form for R :

$$R = \begin{bmatrix} \cos \phi & 0 & -\sin \phi \\ 0 & 1 & 0 \\ \sin \phi & 0 & \cos \phi \end{bmatrix}$$

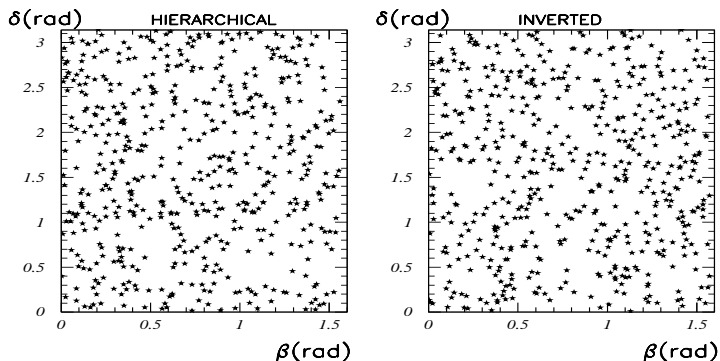
We can write Y_B independently from the low energy phases (with $\phi = \rho + i\omega$) :

$$Y_B \simeq 10^{-10} \left(\frac{M_1}{10^{11} \text{GeV}} \right) \frac{\sin \rho \cos \rho \sinh \omega \cosh \omega}{\sin^2 \rho \cosh^2 \omega + \cos^2 \rho \sinh^2 \omega}$$

Parameter scan

- Random selection of points where the baryon asymmetry is large enough
- Neutrino masses are taken non degenerate
- Dirac phase defined such that $U_{e3} = \sin \theta_{13}^{-i\delta}$
- The baryon asymmetry arises in the decay of N_1 with mass $M_1 = 10^{10}$ GeV

Parameter scan



A large enough baryon asymmetry can be obtained
for any values of the U_{MNS} phases

Leptogenesis in minimal supergravity

Next step :

- Analysis performed in leptogenesis with **minimal supergravity**
- Enhanced flavour violating processes due to RGE running of the sneutrino mass matrix, e.g. :

$$BR(\mu \rightarrow e\gamma) = \frac{\alpha^3 \tan^2 \beta}{G_F^2 m_{SUSY}^8} \left| \frac{1}{8\pi^2} (3M_0^2 + A_0^2) \lambda_{ek} \lambda_{k\mu}^\dagger \log\left(\frac{M_X}{m_{M_k}}\right) \right|^2$$

⇒ Measurable observables : $\lambda\lambda^\dagger = V_L^\dagger D_\lambda^2 V_L$

- Effects on **electric dipole moments**
- Parameter scan with **Markov Chain Monte Carlo**
- **Results soon...**

Conclusions

- The relevant question in discussing “relation” between CP violation in the U_{MNS} matrix :

Is the baryon asymmetry *sensitive* to the U_{MNS} phases ?

- The answer was **NO** for **unflavoured leptogenesis** in the SM seesaw (Branco et al.)
- We argue that the answer does not change also with the inclusion of **flavour effects** in leptogenesis :

For any value of the U_{MNS} phases it is possible to find a point in the space of unmeasurable seesaw parameters such that leptogenesis works