

Neutrino Oscillation Phenomenology

Review of the ISAPP summer school - Valencia, 2008

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Oscillations in vacuum

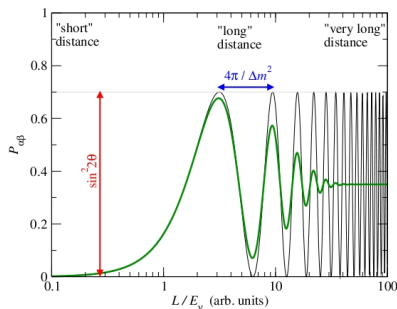
- The oscillation probability in vacuum is given by:

$$P_{\nu_\alpha \rightarrow \nu_\beta}(L, E) = \sum_{kj} U_{\alpha k}^* U_{\beta k} U_{\alpha j} U_{\beta j}^* \exp\left(-i \frac{\Delta m_{kj}^2 L}{2E}\right)$$

- $\Delta m_{kj}^2 \equiv m_k^2 - m_j^2$
→ oscillations are sensitive only to mass-squared differences (not to absolute mass!)
- To observe oscillations one needs:
 - 1 non-trivial mixing $U_{\alpha i}$
 - 2 non-zero mass-squared differences Δm_{kj}^2
 - 3 a suitable value for L/E

Phases of neutrino oscillations

- 1 "short" distance $\Leftrightarrow \Phi \ll 1$:
 - no oscillation can develop
 - $P_{\nu_\alpha \rightarrow \nu_\beta} = \delta_{\alpha\beta}$ because of $\sum_j U_{\alpha j} U_{\beta j}^* = \delta_{\alpha\beta}$
- 2 "long" distance $\Leftrightarrow \Phi \gtrsim \pi/2$:
 - oscillations are observable
- 3 "very long" distance $\Leftrightarrow \Phi \gg 2\pi$:
 - oscillations are averaged out (independent of L and E)



3-flavour oscillation parameters

$$\begin{array}{ccc}
 \Delta m_{31}^2 & & \Delta m_{21}^2 \\
 \left(\begin{array}{ccc} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{array} \right) & \left(\begin{array}{ccc} c_{13} & 0 & e^{-i\delta}s_{13} \\ 0 & 1 & 0 \\ -e^{i\delta}s_{13} & 0 & c_{13} \end{array} \right) & \left(\begin{array}{ccc} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{array} \right) \\
 \text{LBL + atmospheric} & \text{Chooz} & \text{Solar + KamLAND}
 \end{array}$$

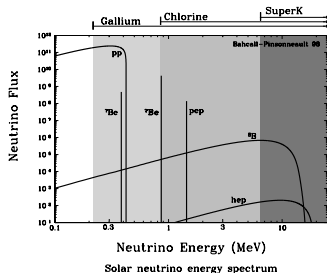
- 3-flavour effects are suppressed because:

- 1 $\theta_{13} \ll 1$
- 2 $\Delta m_{21}^2 \ll \Delta m_{31}^2$

⇒ Dominant oscillations are well described by 2-flavour oscillations

Solar neutrino experiments

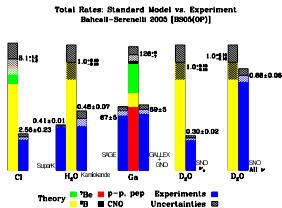
- Homestake, Kamiokande, GALLEX/GNO, SAGE, Super-K, SNO, Borexino
- Thermonuclear fusion reactions in the solar core \Rightarrow source of ν_e



- Accessible oscillation parameters $\rightarrow \theta_{12}, \Delta m_{21}^2$

The solar problem and its solution

- Evidence for an oscillation $\nu_e \rightarrow \nu_{\mu,\tau}$

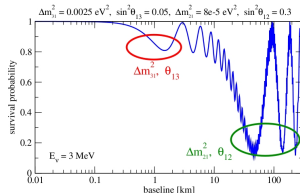


- LMA + MSW:

- $\tan^2 \theta_{12} = 0.45^{+0.09}_{-0.07}$ (SNO)
- $\Delta m_{21}^2 \cos 2\theta_{12} = 2EV \Rightarrow \Delta m_{21}^2 > 0$

Reactor neutrino experiments

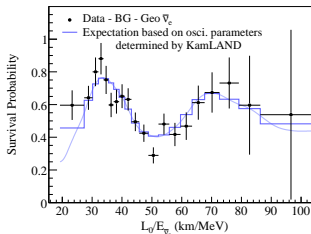
- Bugey, Palo Verde, CHOOZ, KamLAND
- Nuclear fission in reactor cores \Rightarrow source of $\bar{\nu}_e$
- Only disappearance experiment because of $E \approx 4$ MeV



- Accessible oscillation parameters $\rightarrow \theta_{13}$, Δm_{31}^2 or θ_{12} , Δm_{21}^2 depending on the baseline

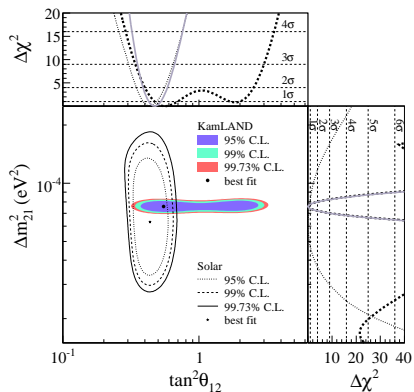
Chooz and KamLAND

- CHOOZ: $\frac{E}{L} \approx \frac{4\text{MeV}}{1\text{km}} \approx 4 \times 10^{-3} eV^2$
 $\Rightarrow \bar{\nu}_e$ disappearance at the "atmospheric" Δm_{31}^2 scale
- KamLAND: $\frac{E}{L} \approx \frac{4\text{MeV}}{175\text{km}} \approx 2 \times 10^{-5} eV^2$
 $\Rightarrow \bar{\nu}_e$ disappearance at the "solar" Δm_{21}^2 scale



- $\Delta m_{21}^2 = 7.59 \pm 0.21 \times 10^{-5} eV^2$

KamLAND vs solar data

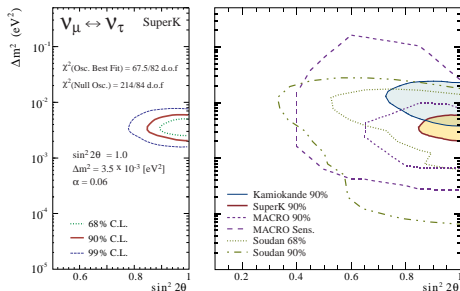


$$\Delta m_{21}^2 = 7.59 \pm 0.21 \times 10^{-5} eV^2 \text{ (KamLAND)}$$

$$\tan^2 \theta_{12} = 0.45^{+0.09}_{-0.07} \text{ (SNO)}$$

Atmospheric neutrino experiments

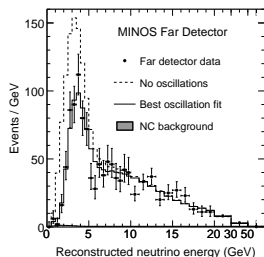
- IMB, Soudan 2, Super-K
- Hadronic showers in the earth's atmosphere $\Rightarrow \nu_{e,\mu} \bar{\nu}_{e,\mu}$
- Accessible oscillation parameters $\rightarrow \theta_{23}, \Delta m_{31}^2$



- $\sin^2 2\theta_{23} = 1.00$ (SK)

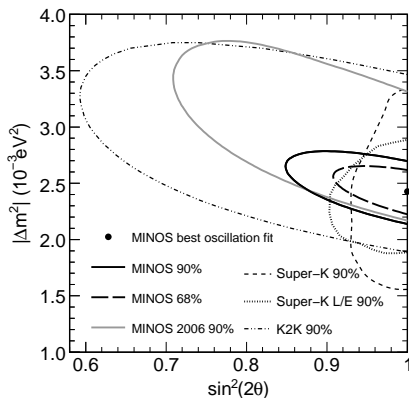
LBL neutrino experiments

- K2K, MINOS \rightarrow disappearance of accelerated ν_μ (or $\bar{\nu}_\mu$) flux
- MINOS: $\frac{E}{L} \approx \frac{3\text{GeV}}{735\text{km}} \approx 4.1 \times 10^{-3} eV^2$
- Accessible oscillation parameters $\rightarrow \theta_{23}, \Delta m_{31}^2$



- $|\Delta m_{31}^2| = 2.43 \pm 0.13 \times 10^{-3} eV^2$ (MINOS)

Super-K + K2K + MINOS



$$|\Delta m_{31}^2| = 2.43 \pm 0.13 \times 10^{-3} eV^2 \text{ (MINOS)}$$

$$\sin^2 2\theta_{23} = 1.00 \text{ (SK)}$$

Summary

- "Known" parameters:

- ① $\Delta m_{21}^2 = 7.59 \pm 0.21 \times 10^{-5} eV^2$ $\sin^2 \theta_{12} = 0.31^{+0.016}_{-0.023}$
- ② $|\Delta m_{31}^2| = 2.43 \pm 0.13 \times 10^{-3} eV^2$ $\sin^2 \theta_{23} = 0.50 \pm 0.063$

- Open questions:

- ① Sign of Δm_{31}^2 ?
 - Look for matter effect in $\nu_e \leftrightarrow \nu_\mu$ oscillations
 - Neutrinoless double beta decay (see Tarek's talk)
- ② Value of $\sin^2 \theta_{13}$?
 - Double Chooz, Daya Bay, T2K (see Rachel's talk)
- ③ Value of the phase δ ?